

# The type D horizons and the Petrov type D spacetimes

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The type D equation with a cosmological constant  $\Lambda$  can be imposed on pairs  $(g, \omega)$ : a metric tensor  $g$  and a differential 1-form  $\omega$  defined on a 2-dimensional manifold  $S$ . If  $S$  is a space-like section of a non-extremal Killing horizon  $H$  contained in a 4 dimensional vacuum spacetime with a cosmological constant  $\Lambda$ , then the equation ensures that the Petrov type of the Weyl tensor at the horizon is  $D$ . The horizon may admit a global section and have the trivial fibration  $S \times \mathbb{R}$  structure or may admit local sections only and have the fibration structure of the Hopf bundle or more general Dirac monopole. We have found all the axially symmetric type D horizons of the trivial fibration structure and of the non-trivial fibration structure  $S_3 \rightarrow S_2$ , respectively. The trivial fibration horizons are all embeddable either in the Kerr/Kerr-dS/Kerr-AdS spacetimes or in the NHG spacetimes. However, the  $S_3$  type D horizons, surprisingly, do not fit the Kerr-NUT spacetimes. Still, via the black hole holograph construction, every axisymmetric type D horizon must be embeddable in a Petrov type D vacuum solution to Einstein's equation with a cosmological (possibly vanishing) constant. All that puzzle will be presented and discussed in our talk.

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